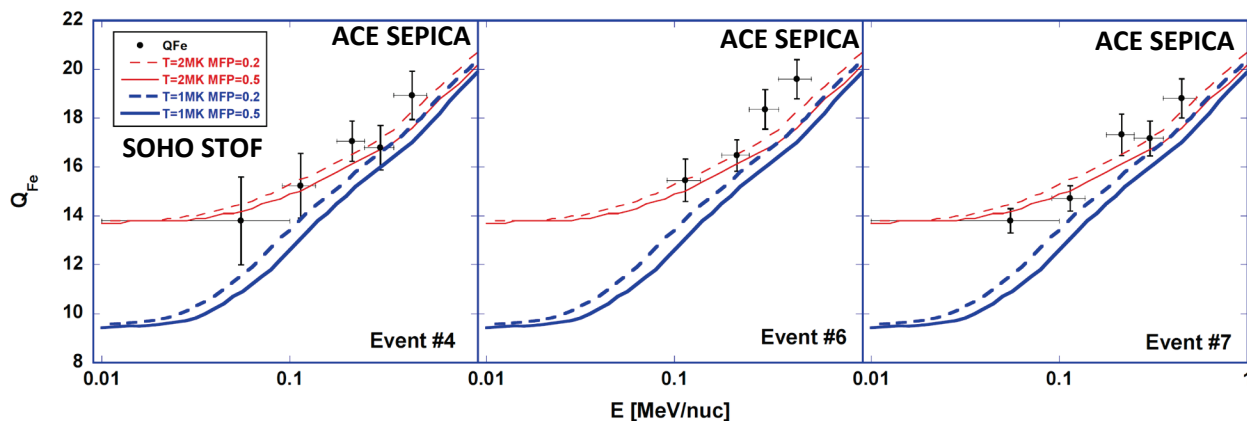


ACE News #168 – April 23, 2014

Impulsive Events with High Charge State Fe at Low Energies: Occasional High Temperature Source Material!



Mean Fe charge states for three impulsive energetic particle events as a function of energy together with model curves, including acceleration, stripping, and adiabatic cooling for source temperatures of 1 MK (blue) and 2 MK (red), and transport mean free paths $\lambda = 0.5$ AU (solid) and $\lambda = 0.2$ AU (dashed) (Kartavykh et al., *ApJ*, 671, 947, 2007). Typical coronal temperatures as derived from ionic charge states in the solar wind are about 1 MK ([ACE News #146](#)).

Earlier studies have shown that impulsive Solar Energetic Particle (SEP) events emit Fe ions with $Q_{Fe} < 14$ at the lowest energies ($E \leq 0.1$ MeV nuc⁻¹) that are consistent with typical corona source material ([ACE News #107](#), DiFabio et al., *ApJ*, 687, 623, 2008). However, the occasional observation of Fe with $Q_{Fe} > 16$ in solar wind associated with active regions ([ACE News #52](#), Lepri et al., *JGR*, 106, 29231, 2001) led to a search for acceleration of high charge state Fe in all SEP events observed with ACE SEPICA.

A total of nine SEP events with $Q_{Fe} \geq 14$ across the entire SEPICA energy range (0.08 - 0.54 MeV/nuc) were identified (Guo et al., *ApJ*, 785, 26, 2014). Five of these events were CME related events with acceleration of high charge state material, in some cases mixed with the acceleration of a separate component of more normal coronal material. Four of these events are impulsive events with the clear signature of ion velocity dispersion. In comparison with a model that involves acceleration out of a source at equilibrium temperature, further stripping during the acceleration, and adiabatic cooling during transport to 1 AU (shown for three events in the Figure above), all four events are consistent with a source temperature $T > 2$ MK (Kartavykh et al., *ApJ*, 671, 947, 2007) and thus the presence of hot solar material in these impulsive events.

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